



International
Centre for
Radio
Astronomy
Research



Curtin University



THE UNIVERSITY OF
WESTERN
AUSTRALIA



Government of Western Australia
Department of the Premier and Cabinet
Office of Science

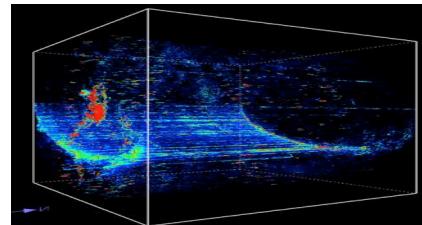
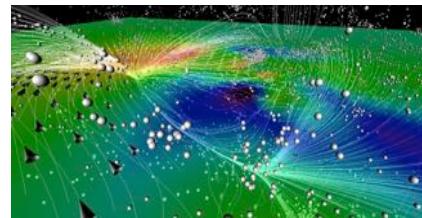
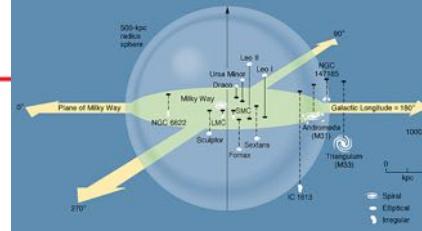


Outline

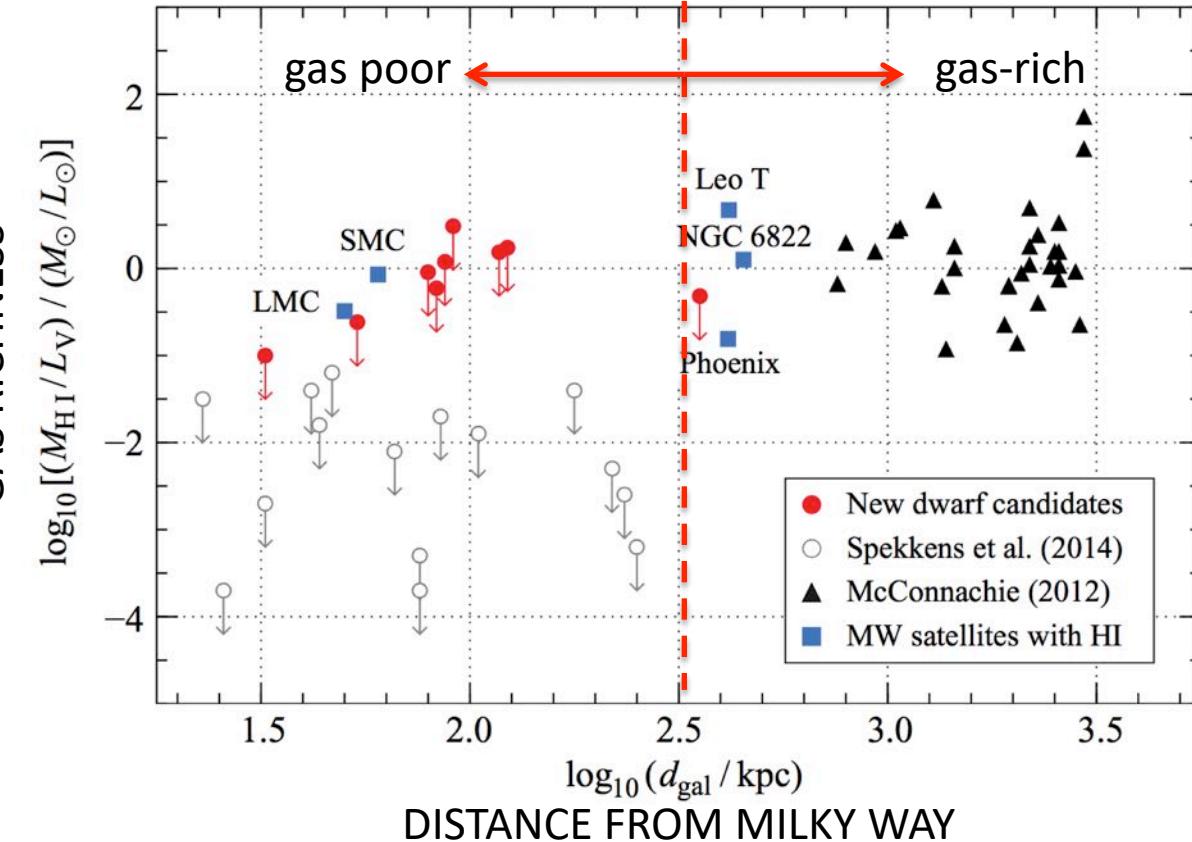
- HI surveys – science goals
- Wallaby
- Cosmic flows
- Northern counterparts
- TAIPAN synergies

Wallaby goals

- Local Group/Volume
- Understanding Galaxies
- Cosmology
- Legacy



Environmental suppression



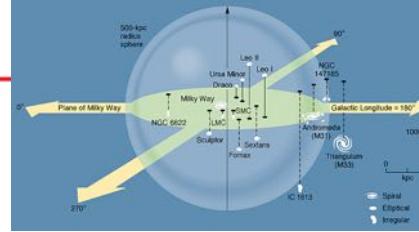
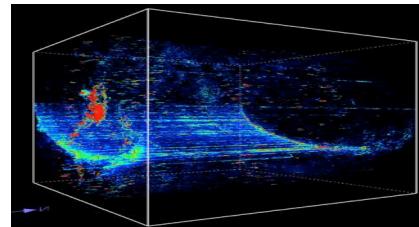
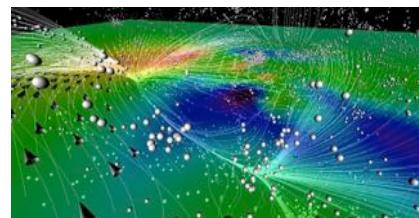
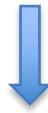
No transition (gas-bearing) galaxies within 300 kpc of Milky Way:

- stripping or ionization?

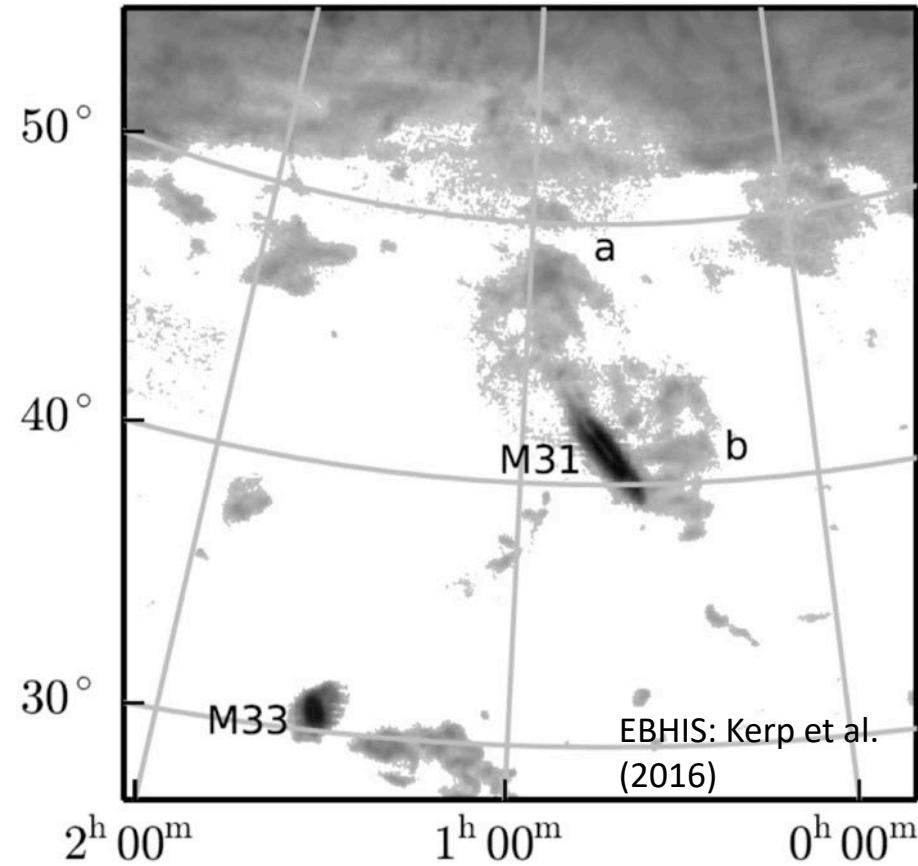
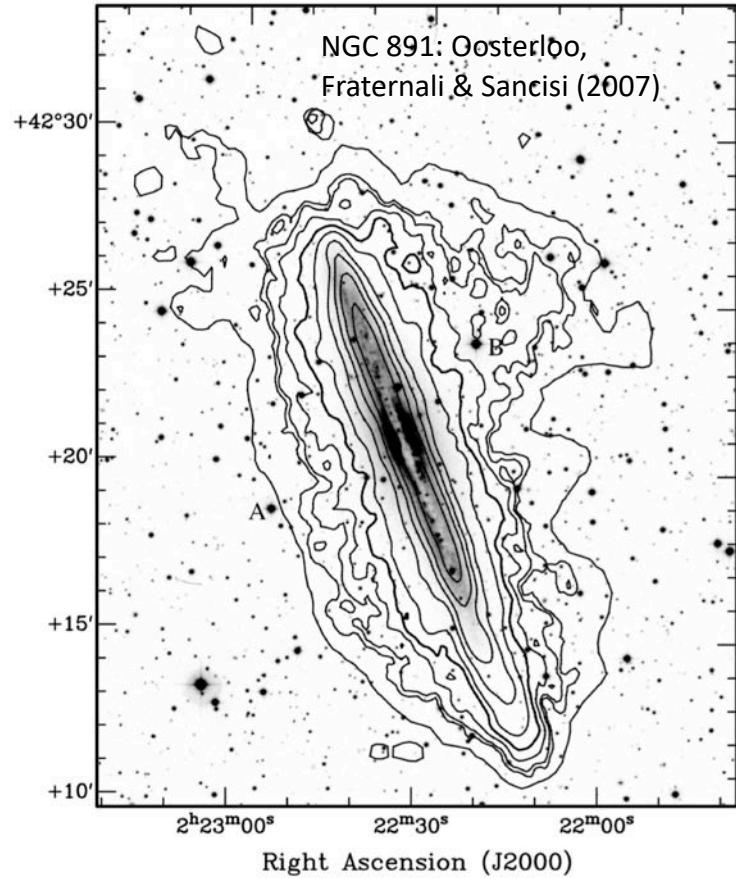
(Westmeier et al. 2015;
Spekkens et al. 2014)

ASKAP Wallaby

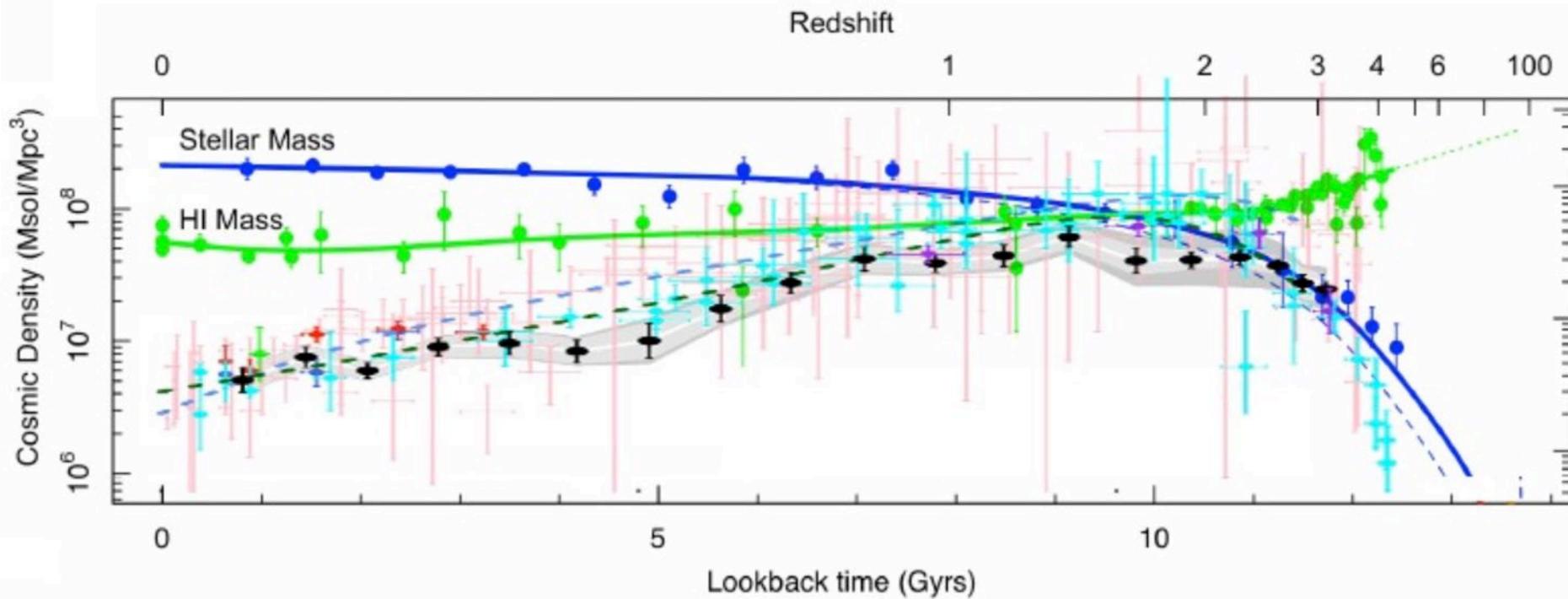
- Local Group/Volume
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Gas accretion onto galaxies



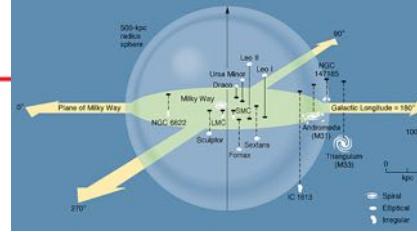
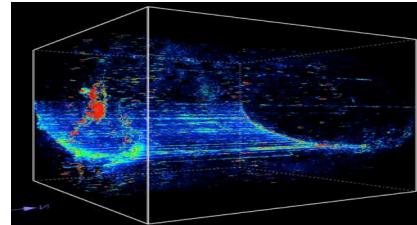
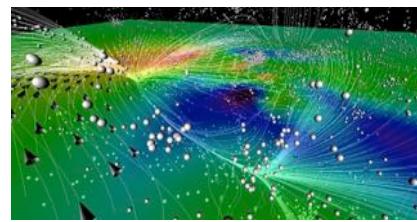
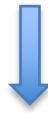
Galaxy evolution



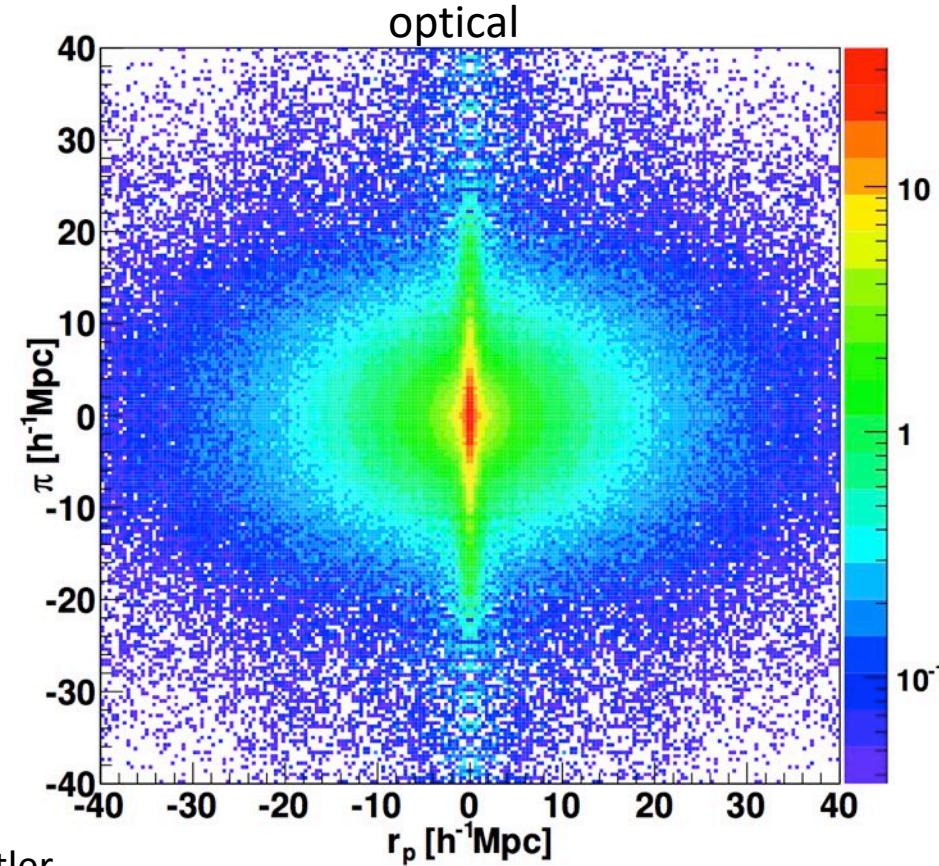
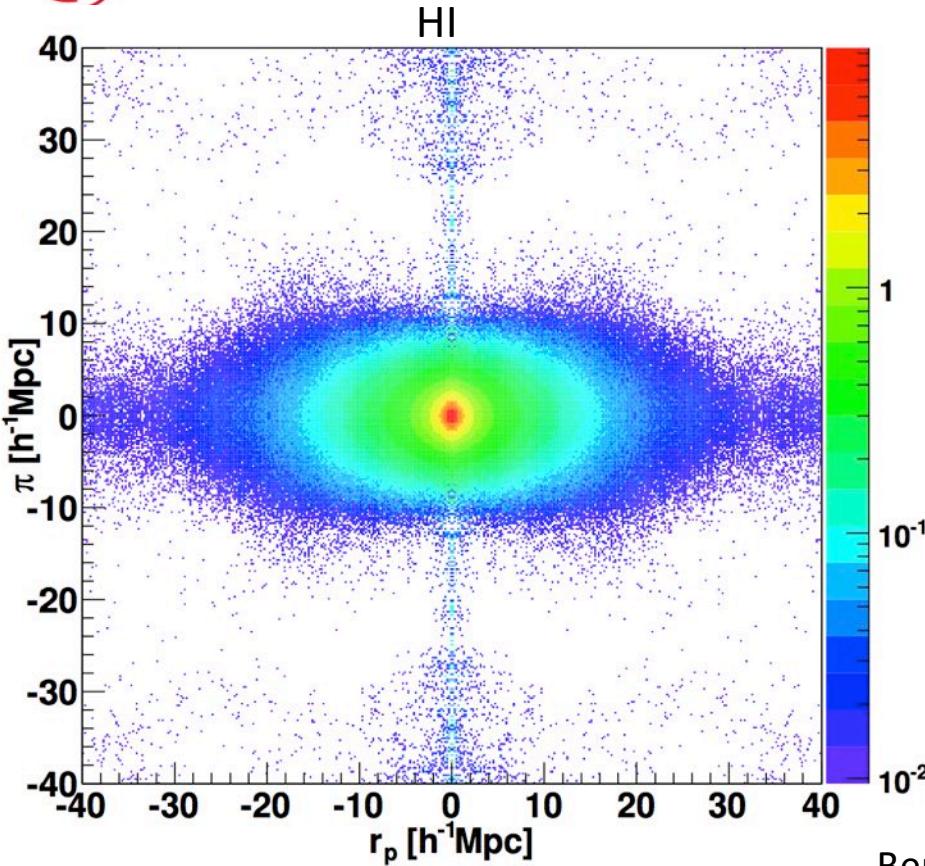
Driver et al. (2018)

ASKAP Wallaby

- Local Group/Volume
- Understanding Galaxies
- Cosmology
- Legacy



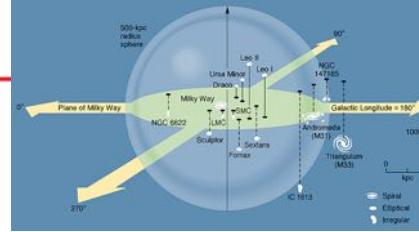
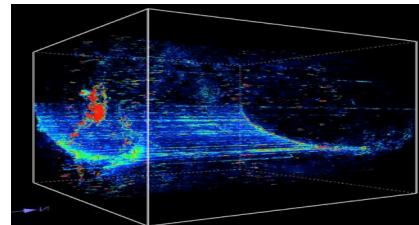
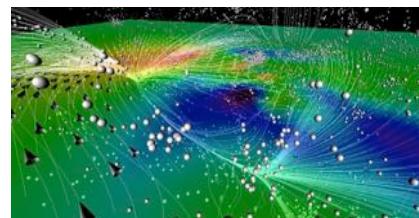
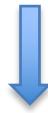
Redshift-space distortions



Beutler

ASKAP Wallaby

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- Cosmology
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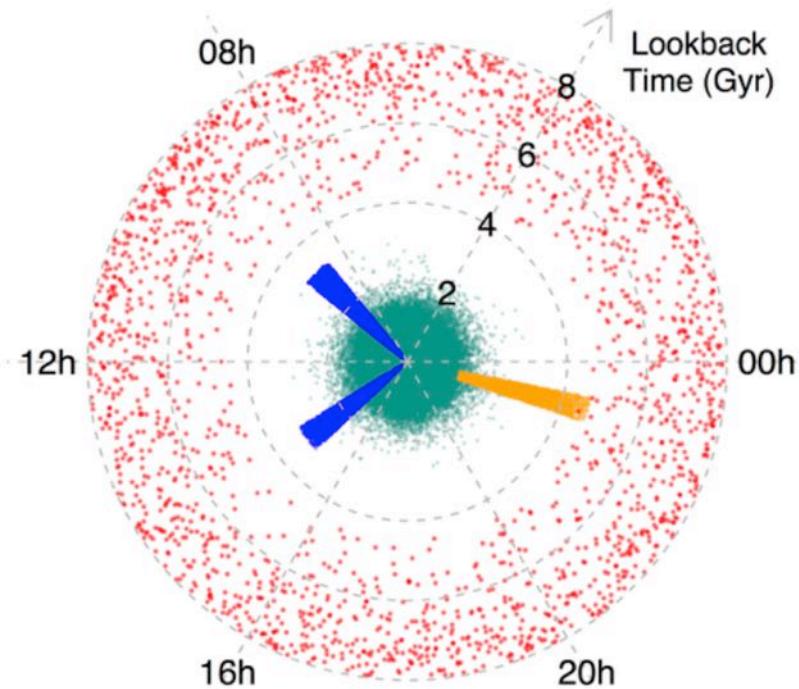
Comparison with other legacy surveys



Area	3π	2.8π	0.7π
Redshift	0 - 0.26	0 - 0.04	0 - 0.06
rms _{20 km/s}	0.7 mJy/beam	13 mJy/beam	1.5 mJy/beam
Res	30 arcsec, 4 km/s	15 arcmin, 18 km/s	3.5 arcmin, 5 km/s
N _{det}	500,000	5,000	30,000

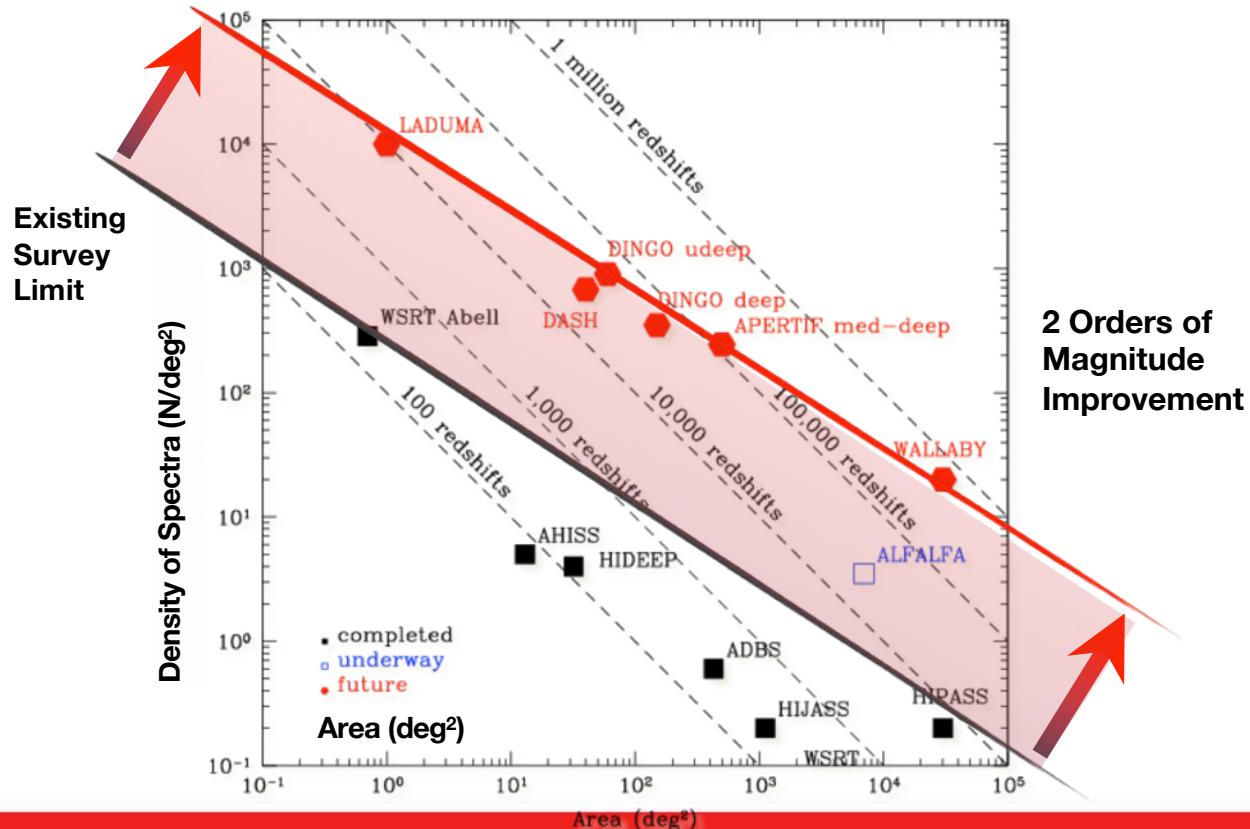
Wallaby = Widefield ASKAP L-band Legacy All-sky Blind surveY

ASKAP HI surveys

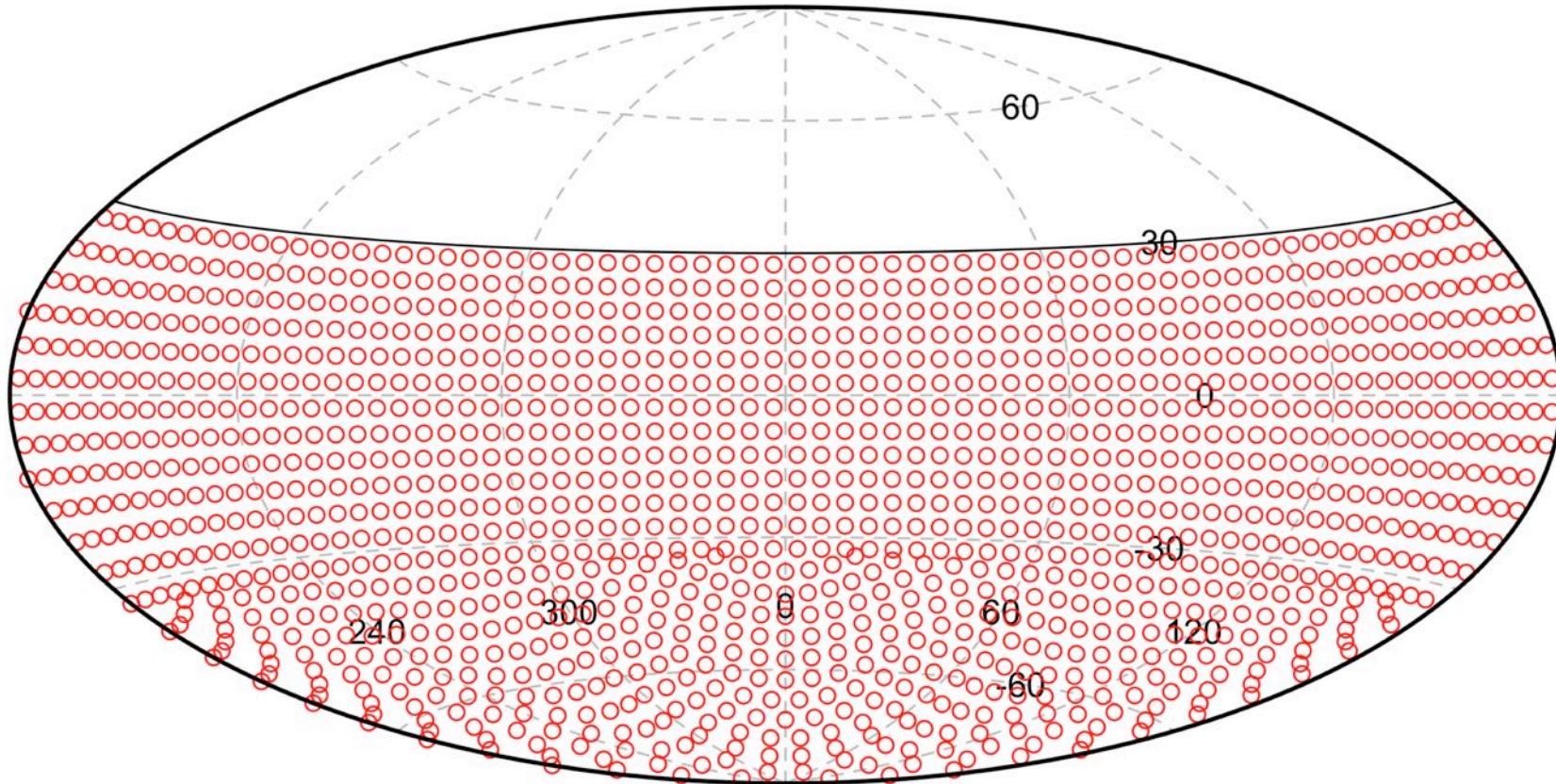


WALLABY DINGO-Deep DINGO-UDeep FLASH

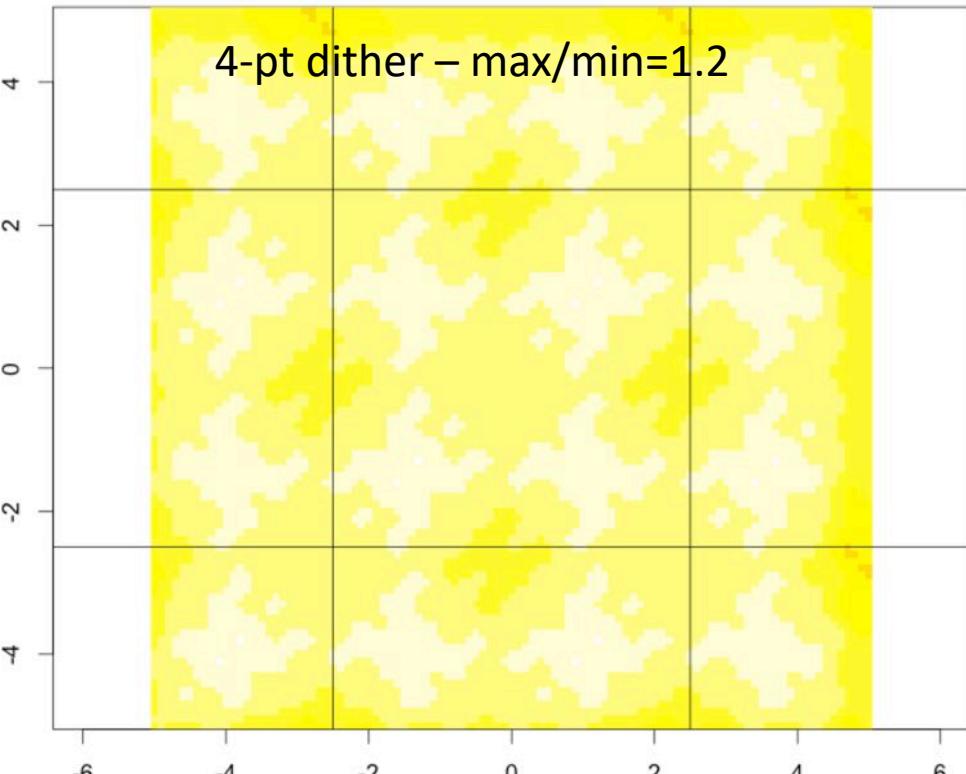
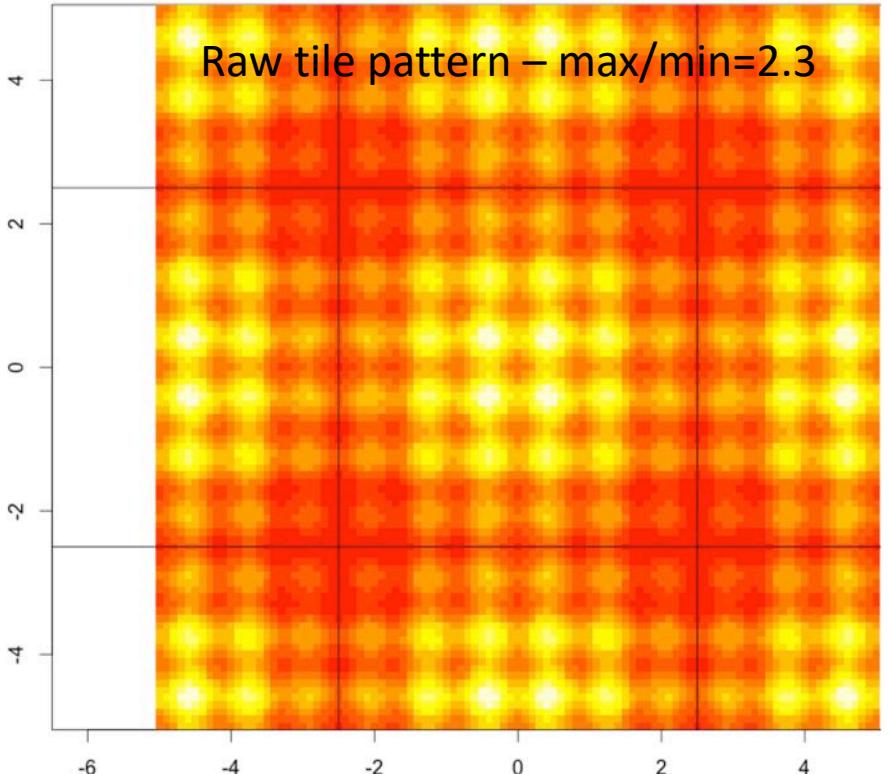
SKA pathfinder HI surveys:



Spherical cap tiling (Robotham)



Sensitivity images (Robotham)



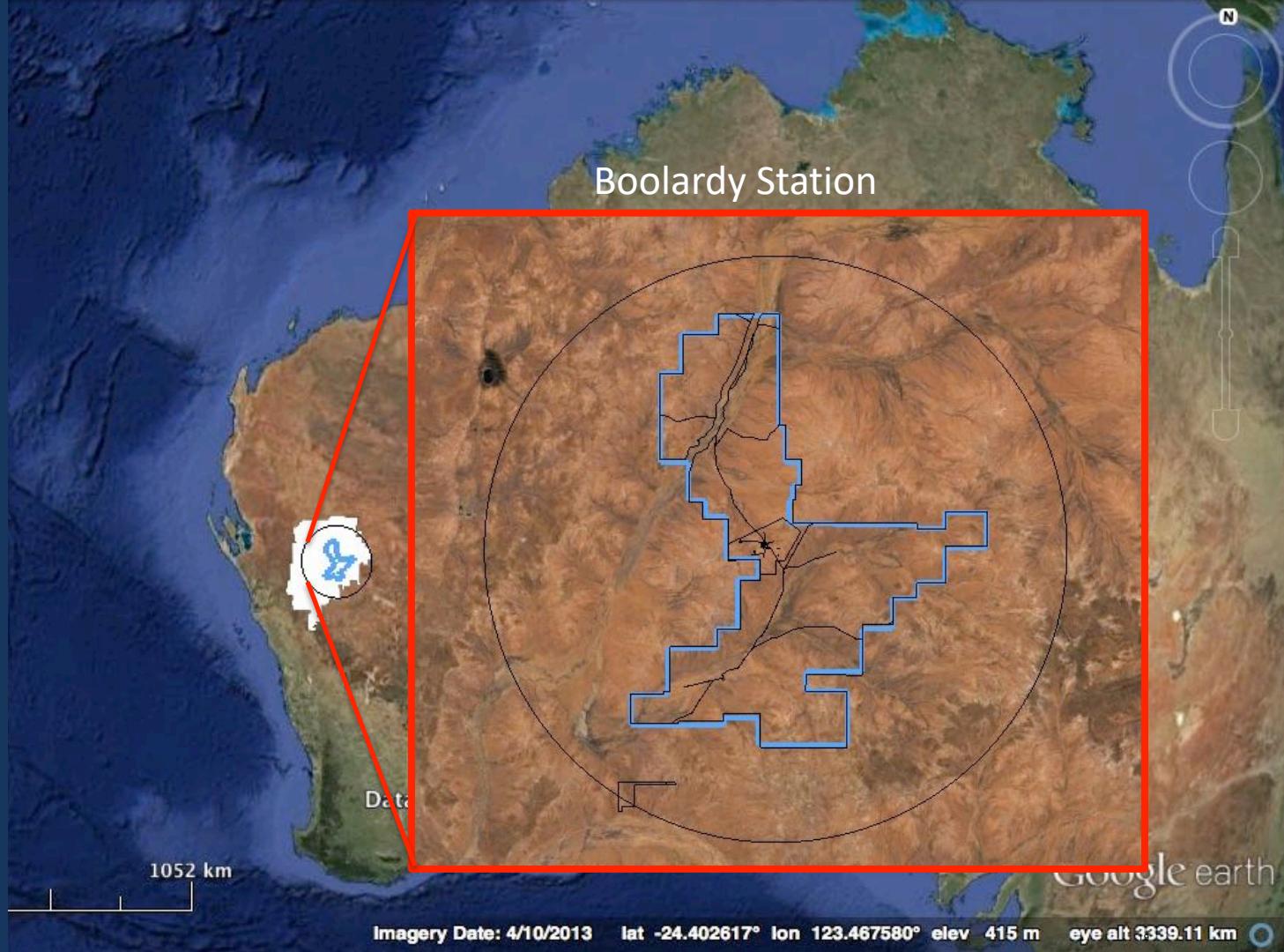
Australian SKA Pathfinder (ASKAP)

- Radio interferometer – $36 \times 12\text{-m}$ dishes
- CSIRO Phased Array Feed (PAF) technology: 36 beams
- Frequency range 700 to 1800 MHz
- Instantaneous bandwidth 300 MHz
- Maximum baseline 6 km
- SKA precursor on a radio-quiet site
- Wallaby (HI) and EMU (continuum) are the two key survey science projects



Imagery Date: 4/10/2013 lat -24.402617° lon 123.467580° elev 415 m eye alt 3339.11 km



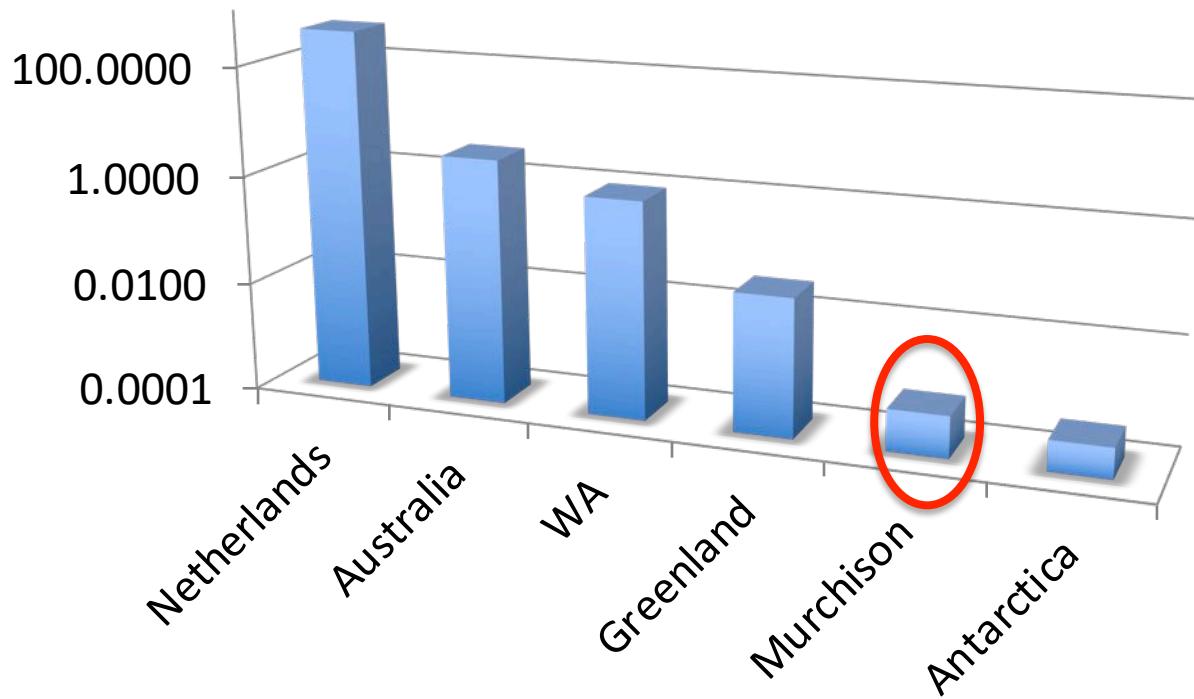


Boolardy Station

Google earth

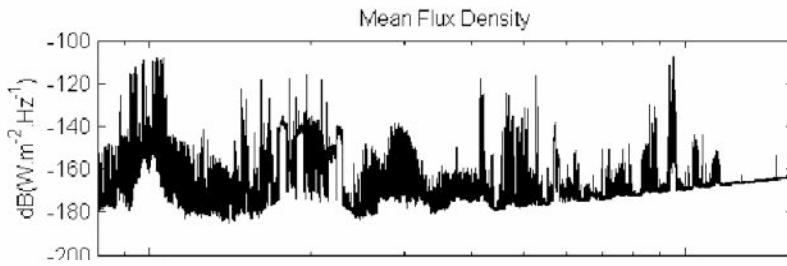
Imagery Date: 4/10/2013 lat -24.402617° lon 123.467580° elev 415 m eye alt 3339.11 km

Population density (people/km²)

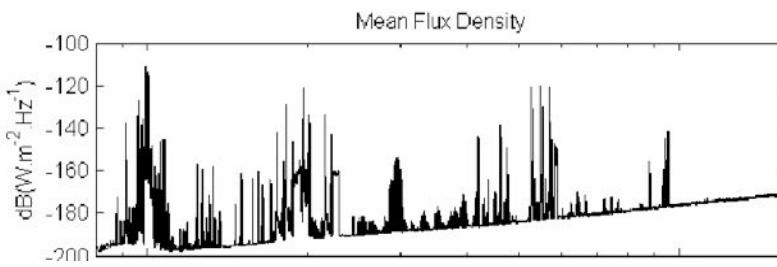


A radio-quiet site

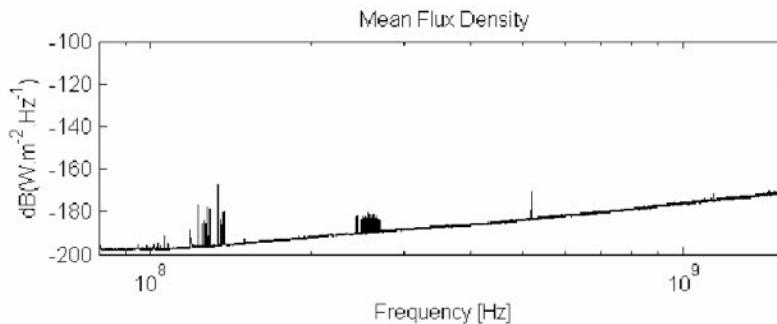
Sydney:
Population 5M



Narrabri (ATCA):
Population 6,000



Murchison, WA
Population 160



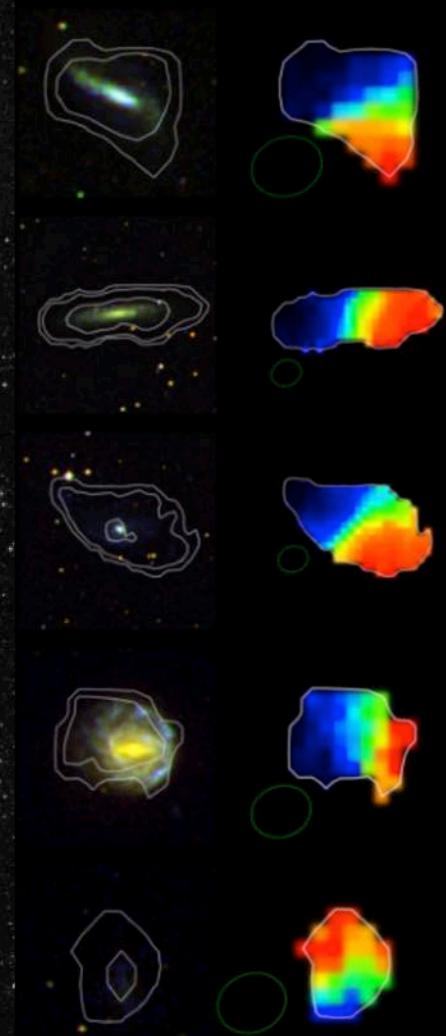
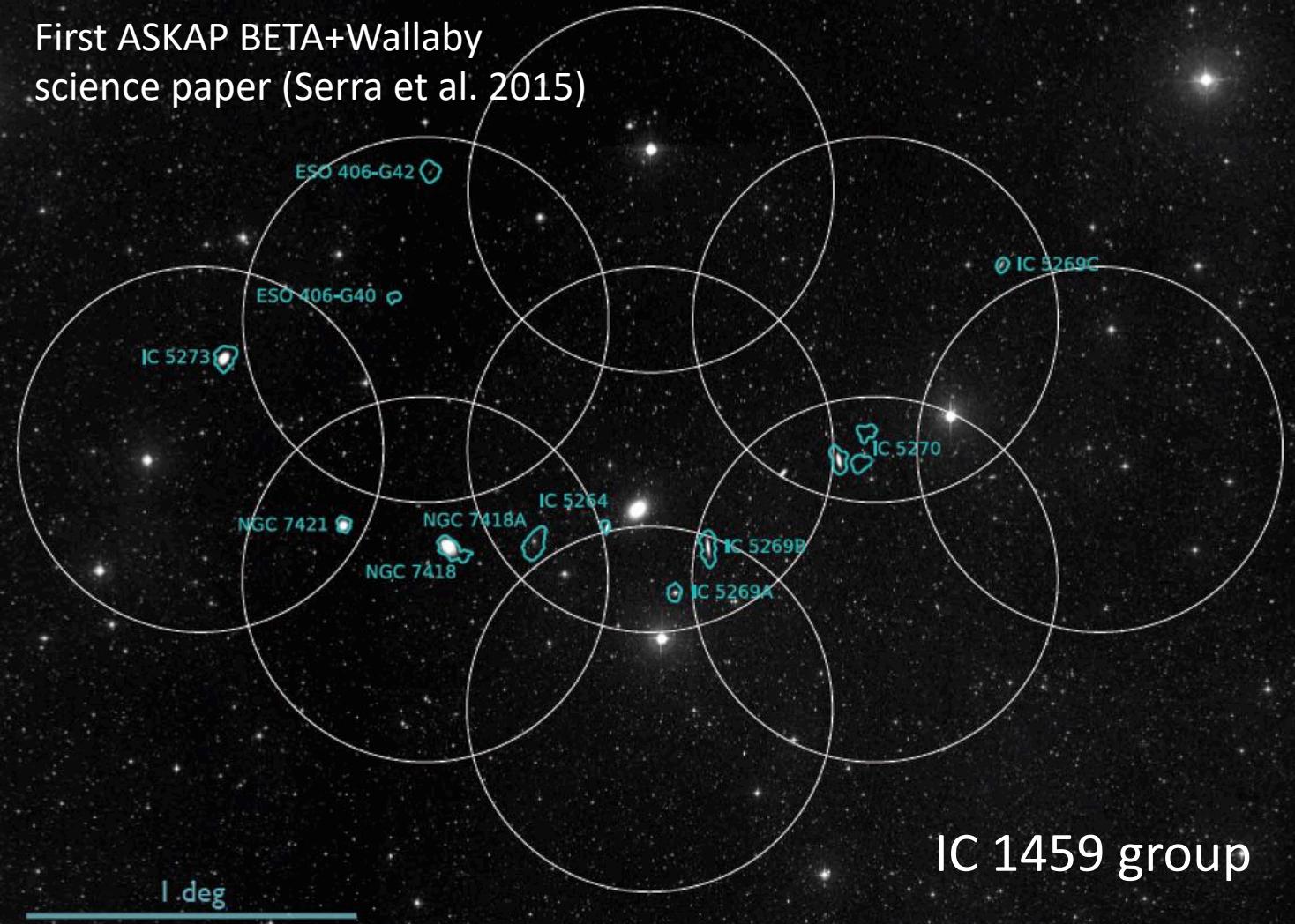




Wallaby timeline

- Sept 2014: commissioning science with BETA (6 antennas)
- Oct 2016: early science commences (12 antennas, 48 MHz)
- Oct 2017: early science concludes (16 antennas, 240 MHz, 650 hrs)
- Dec 2018: Full ASKAP functionality (36 antennas)
- Mar 2019: pilot survey commences (400 hrs)
- Late 2020: Pawsey supercomputer upgrade

First ASKAP BETA+Wallaby science paper (Serra et al. 2015)



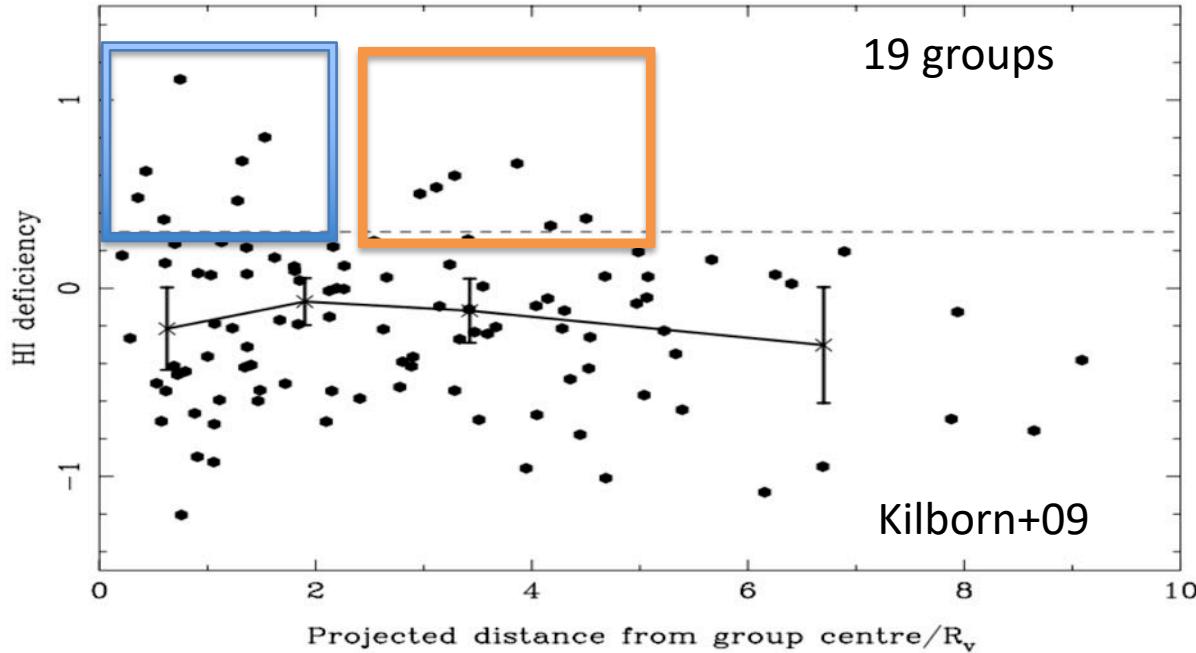


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Early Science (ES) focus: environment

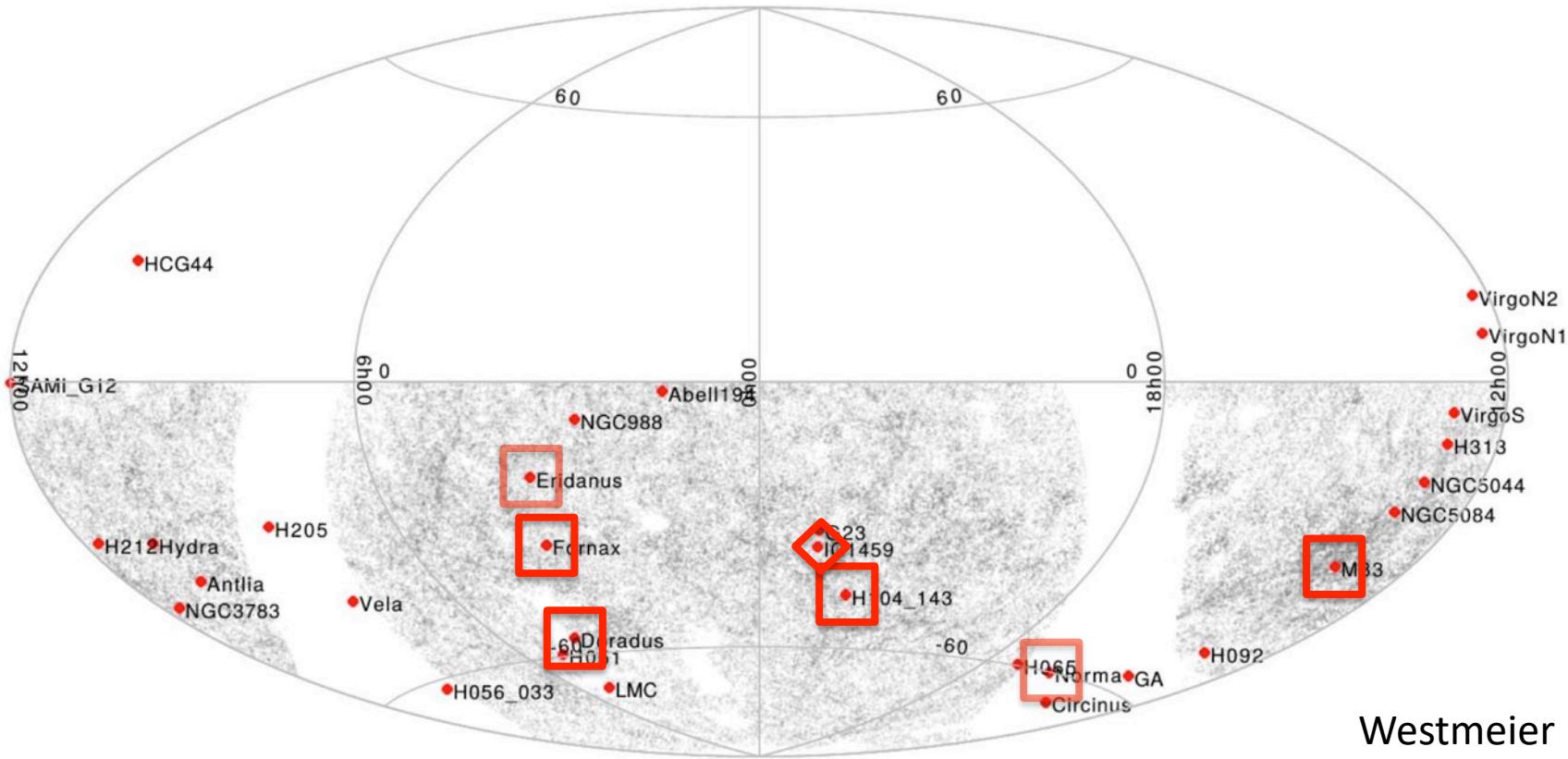
When did galaxies lose their gas?



Preprocessing **in**
and **outside**
groups

Wallaby SWG3
pre-processing
workshop –
Swinburne March
8-10 (Kilborn and
Wong)

WALLABY ES fields



Early Science progress



WALLABY Early Science * Observations with ASKAP-12

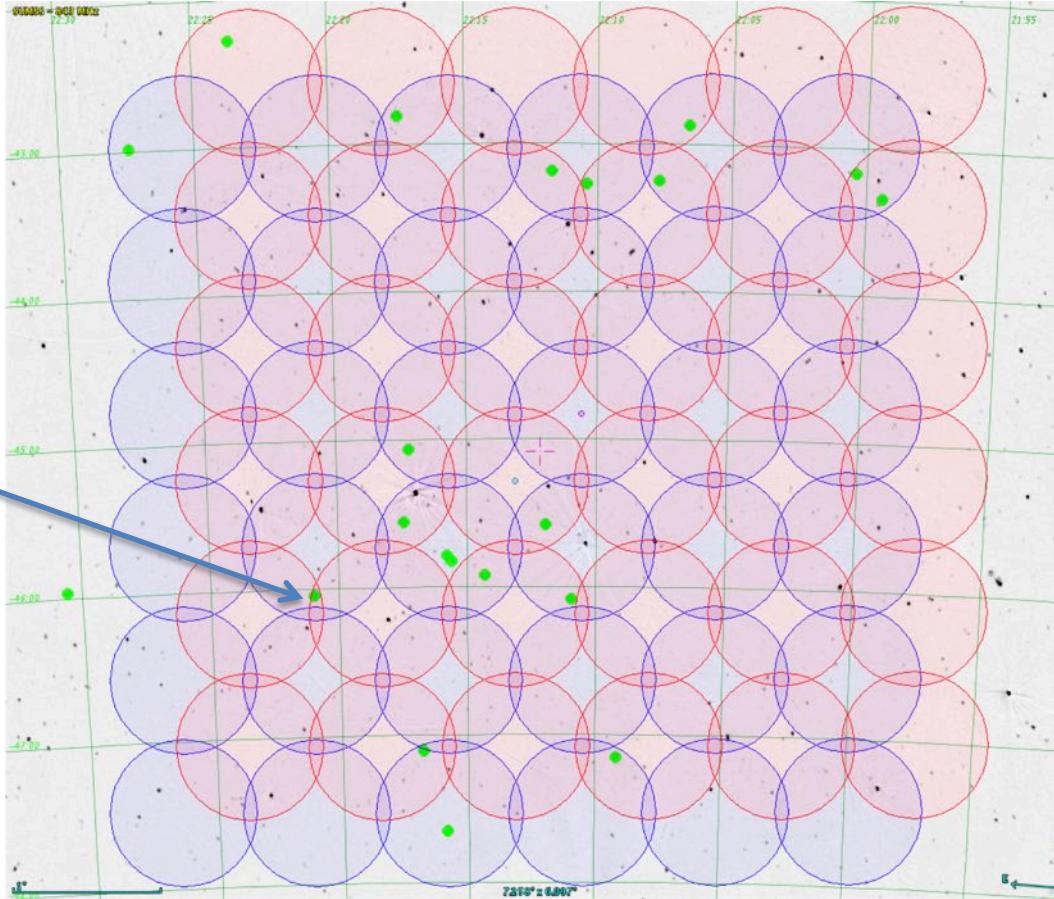
No	Field name	Date	Bandwidth [MHz]	Antennas	Flagged [%]	Total time [hours]
Field 1	NGC 7232 Group	Aug to Sep-16	48	10 - 12	16 - 43	43.6
"	"	Oct-16	48	10 - 12	11 - 28	136.8
Field 2	Fornax Cluster	Oct to Dec-16	48/144/168	10 - 12	20 - 34	59.3
"	"	Dec-16	192	10 - 12	11 - 30	163.9
Field 3	Dorado Group	Dec-16/Jan-17	192	9 - 10	tbd	72.9
"	"	Sep-17	192/240	12	tbd	64
Field 4	M 83 Group	Dec-16/Jan-17	192	9 - 10	tbd	80.8
"	"	Sep-17	192/240	12	tbd	31.5

ES data is virtually identical to the Wallaby main survey data in resolution and sensitivity



ES1 coverage

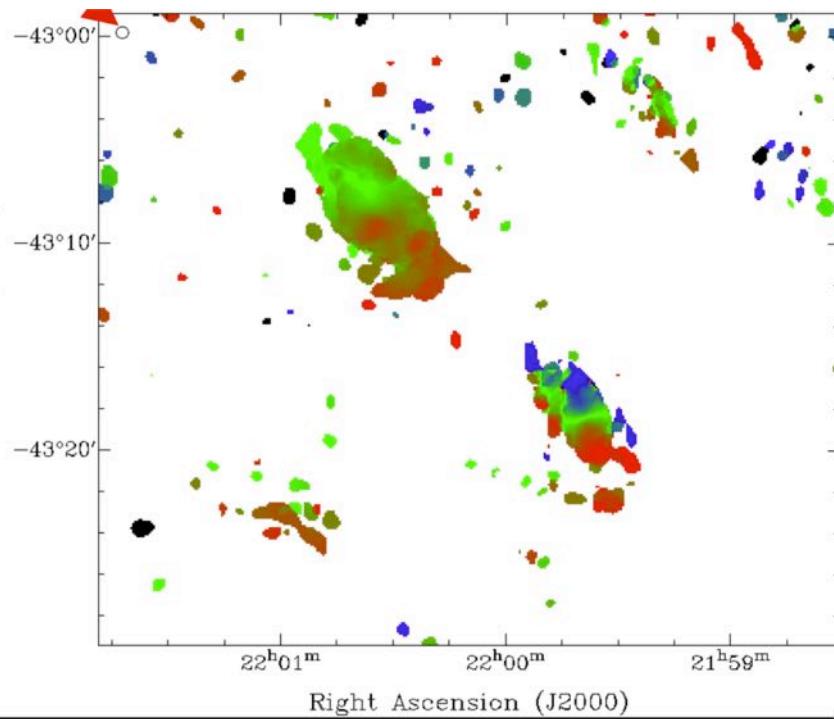
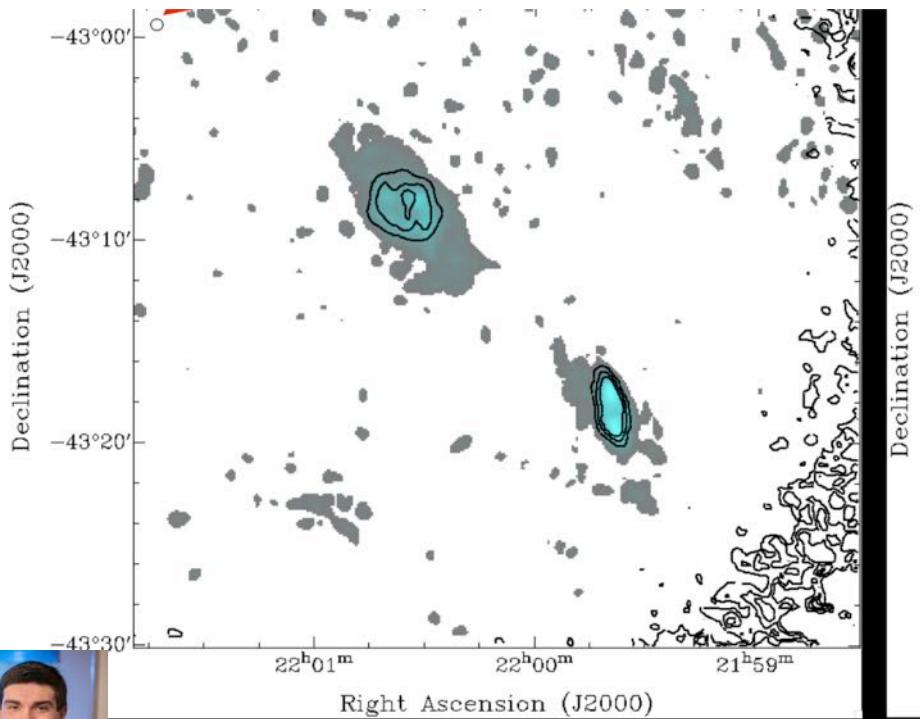
IC 5201



ES1

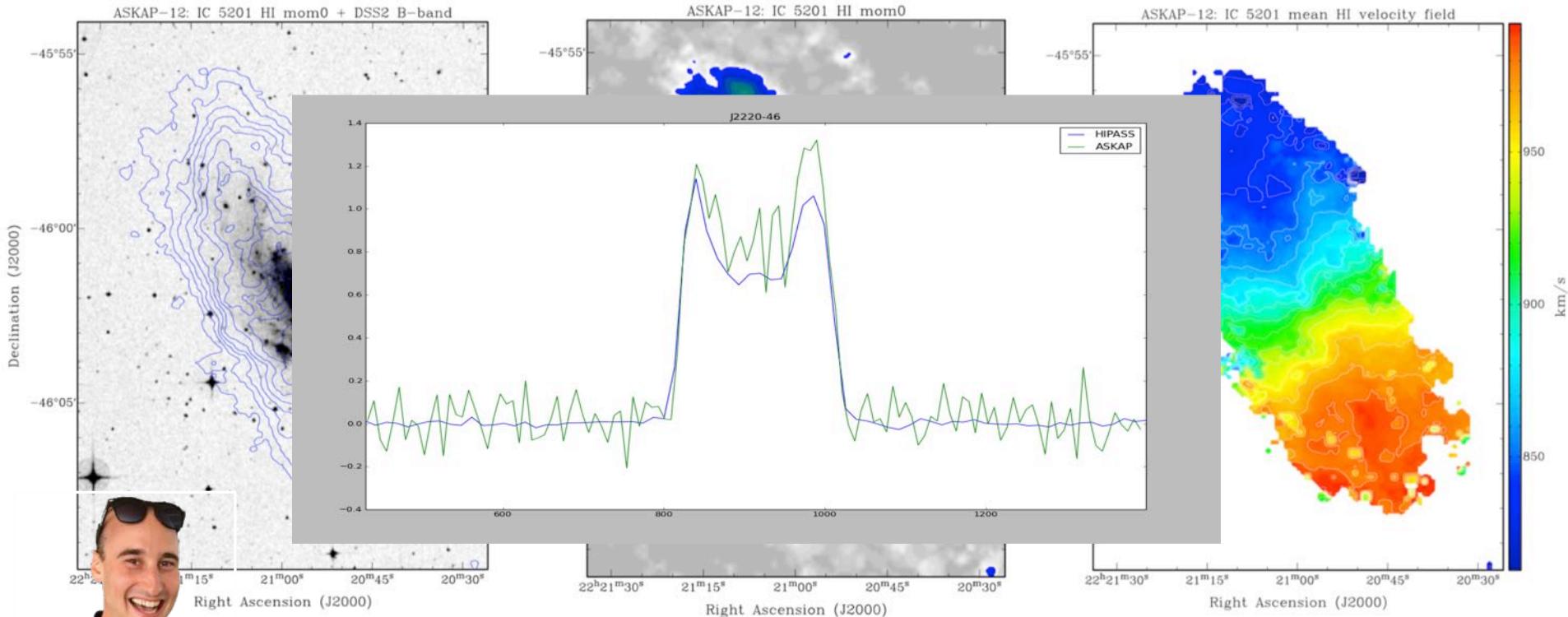
NGC 7232 field
and beam pattern
on coloured DSS

NGC 7162/A in the NGC 7232 group



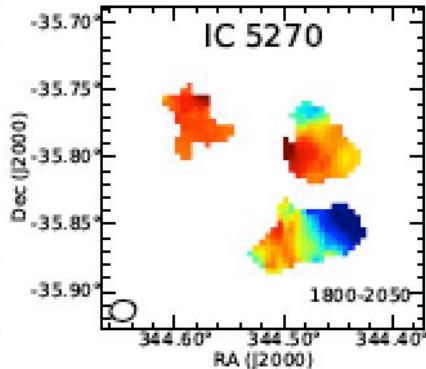
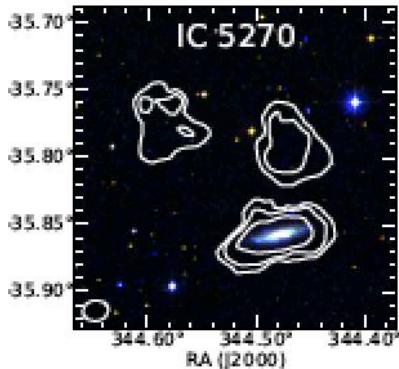
Reynolds et al. (submitted)

IC 5201 in the NGC 7232 group

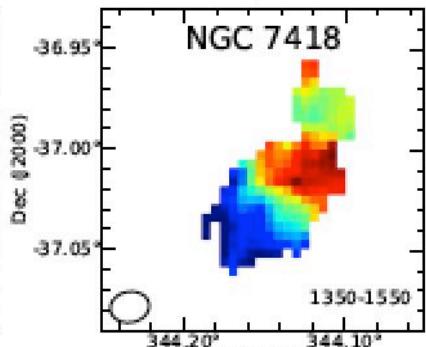
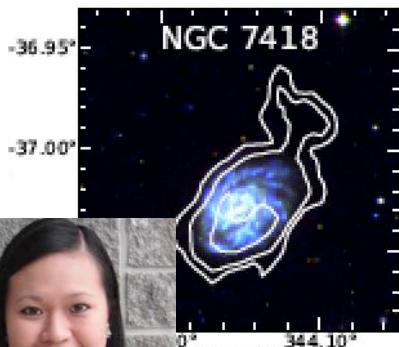


Kleiner et al. (to be submitted)

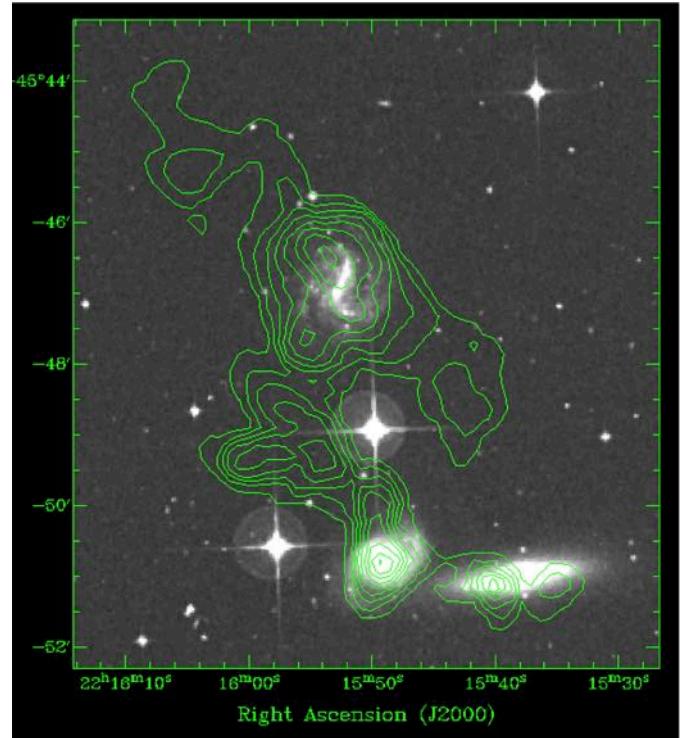
Tidal dwarf galaxies



←
BETA
HI clouds



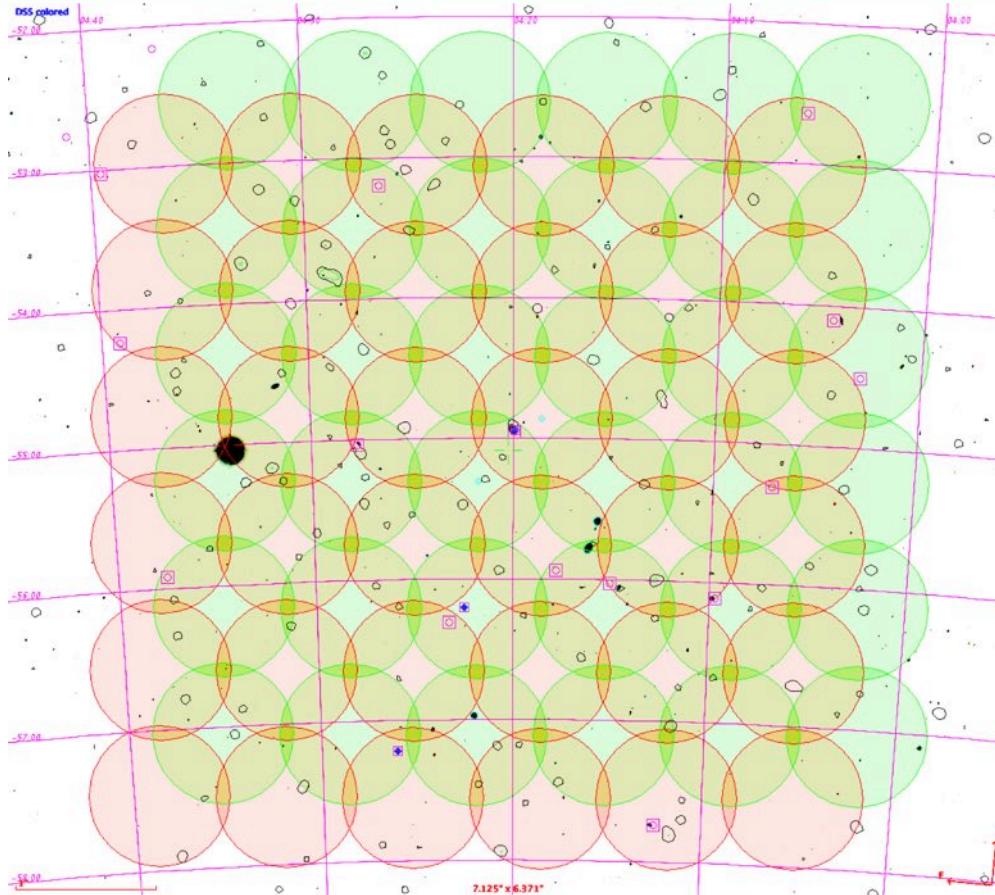
→
ASKAP-12 HI
clouds



Lee-Waddell et al. (submitted)



ES3 coverage

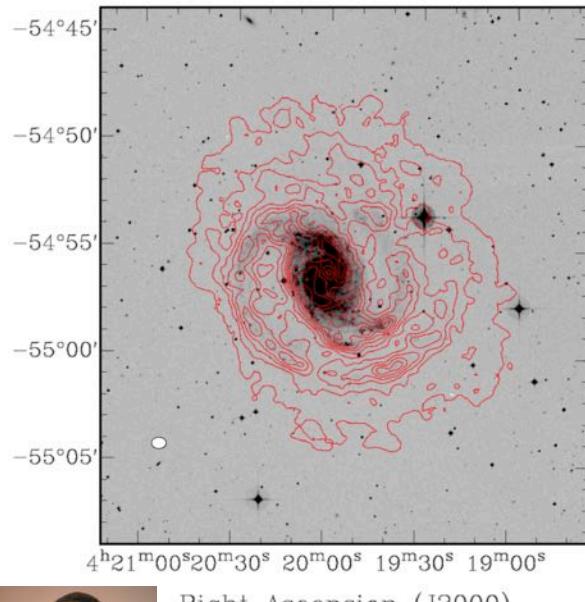


6 deg

ES3

Dorado field and beam pattern on coloured DSS

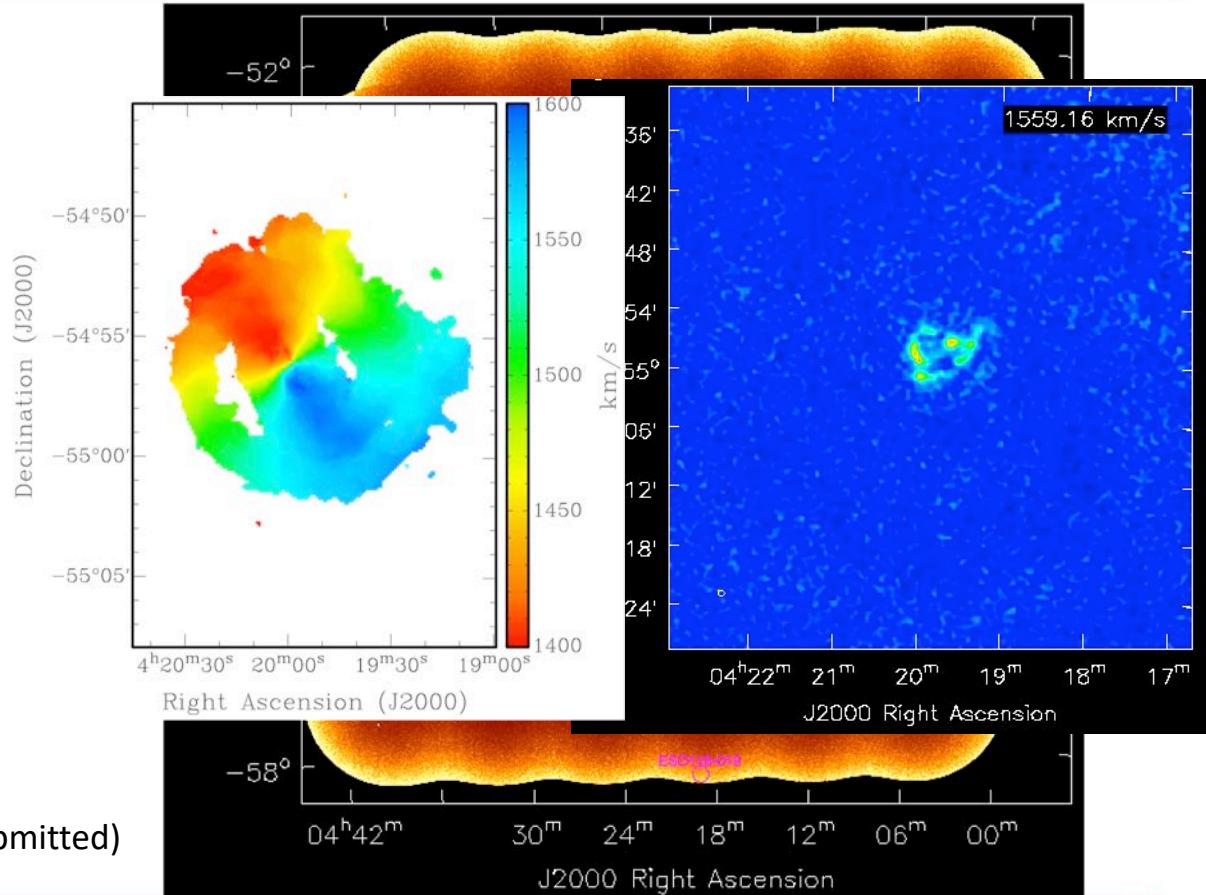
Dorado



Right Ascension (J2000)

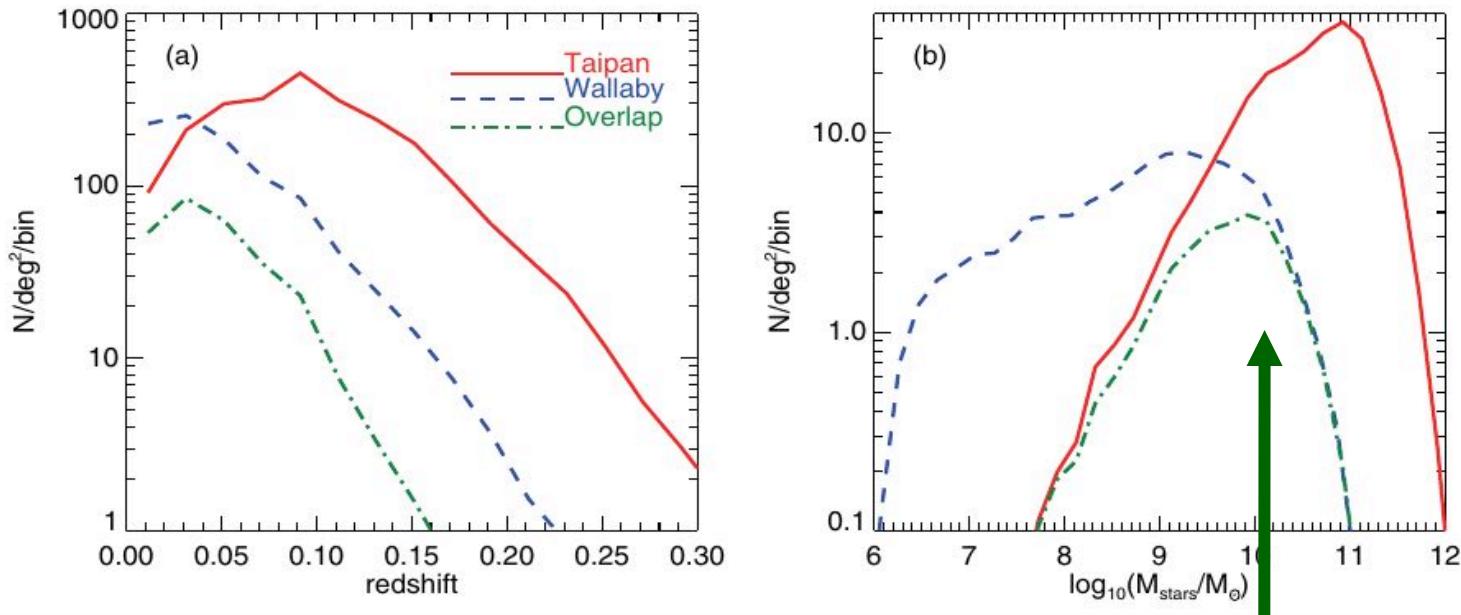


Agali, Rhee et al (to be submitted)



Taipan/WALLABY synergies

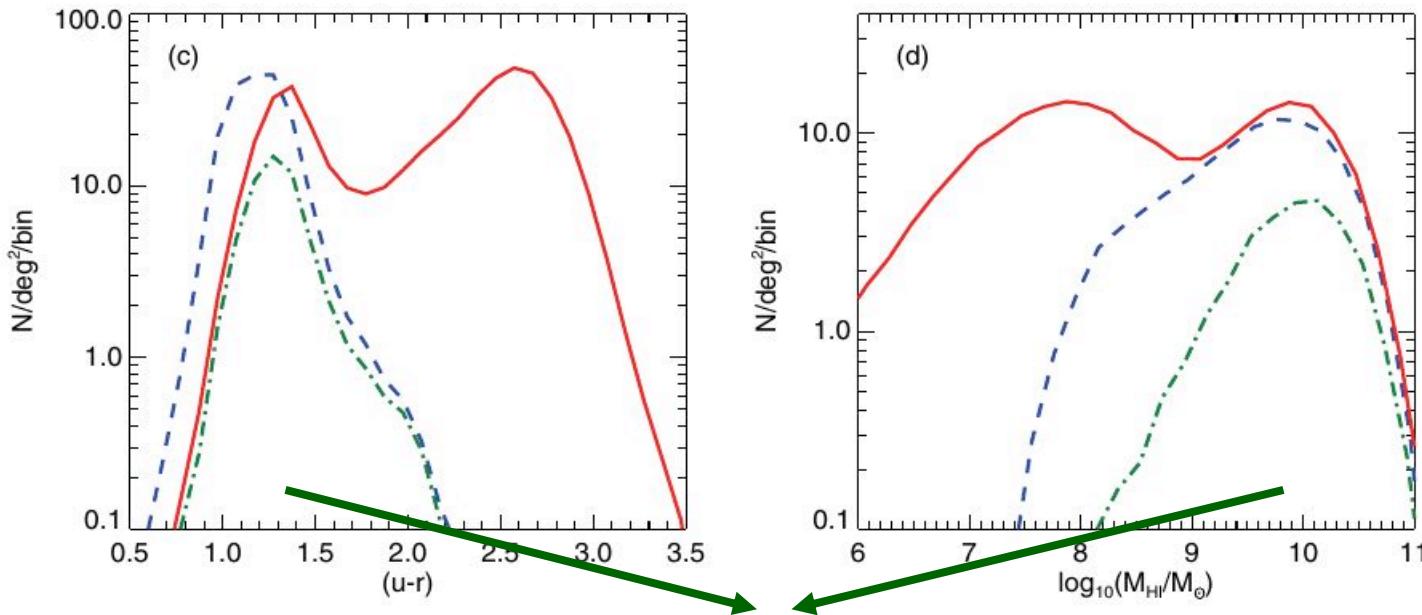
Da Cunha et al. (2017): Taipan white paper – Predictions based on GALFORM model of Lagos et al. (2012) using Mill I and II



Overlap dynamic range at intermediate stellar masses:
→ high mass galaxies are gas poor
→ low mass galaxies too faint for optical survey

Taipan/WALLABY synergies

Da Cunha et al. (2017): Taipan white paper – Predictions based on GALFORM model of Lagos et al. (2012) using Mill I and II



Overlap region:
→ blue cloud and green valley
→ high HI masses

Summary

- Wallaby science has started (4+2 ES fields - same area, resolution and sensitivity as ASKAP-36)
- Initial science focus: pre-processing in group environment
- Wallaby is zero-redshift, all-sky ‘anchor’ for high-z studies – most O/IR synergy with TAIPAN, SkyMapper, VISTA, VST, LSST
- Expected completion 2020/21